Supplementary Materials: Performance of Solar-Induced Chlorophyll Fluorescence in Estimating Water-Use Efficiency in a Temperate Forest

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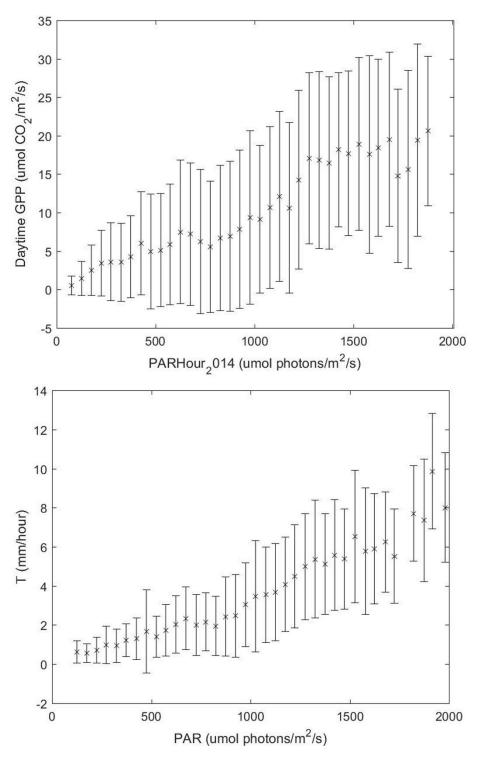


Figure S1. The responses of hourly daytime (a) gross primary production (GPP, umol CO₂ m/s) and (b) transpiration (T, mm/hour) to incident photosynthetically active radiation (PAR, µmol photons/m²/s) during the growing season in 2014. The hourly ET time series was bin-averaged into 50 µmol photons/m²/s increment for PAR. Bars represent standard errors.

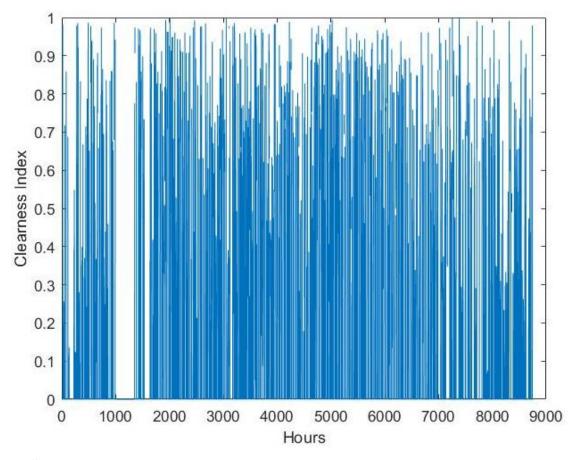


Figure S2. The hourly clearness index (Gu et al., 1999) at the Harvard site in 2014.

Table S1. The correlation coefficient (R^2) of the training and testing groups in predicting daily water use efficiency (WUE), intrinsic water use efficiency (WUE_i), inherent water use efficiency (IWUE) and underlying water use efficiency (uWUE) by using the linear regression analysis (LR) and Gaussian processes regression (GPR), respectively. The training group has 60% of all daily measurements and the test group has the rest. The R^2 estimated from the testing group are provided in the parentheses.

| | WUE | | WUE_{i} | | IWUE | | $\mathbf{u}\mathbf{W}\mathbf{U}\mathbf{E}$ | |
|---|-------------|-------------|-------------|-------------|-------------|-------------|--|-------------|
| | LR | GPR | LR | GPR | LR | GPR | LR | GPR |
| SIF ₆₈₇ | 0.00 (0.00) | 0.01 (0.00) | 0.02 (0.01) | 0.02 (0.01) | 0.20 (0.16) | 0.38 (0.33) | 0.09 (0.06) | 0.34 (0.32) |
| SIF ₇₂₀ | 0.13 (0.13) | 0.39 (0.35) | 0.31 (0.26) | 0.30 (0.27) | 0.48 (0.44) | 0.53 (0.47) | 0.42 (0.38) | 0.41 (0.39) |
| SIF ₇₆₁ | 0.14 (0.11) | 0.40 (0.38) | 0.39 (0.35) | 0.48 (0.43) | 0.58 (0.55) | 0.58 (0.51) | 0.43 (0.39) | 0.52 (0.48) |
| SIF ₇₂₀ , SIF ₇₆₁ | 0.15 (0.11) | 0.50 (0.50) | 0.48 (0.43) | 0.48 (0.41) | 0.60 (0.56) | 0.60 (0.56) | 0.50 (0.47) | 0.58 (0.54) |
| SIF ₆₈₇ , SIF ₇₂₀ , SIF ₇₆₁ | 0.21 (0.18) | 0.52 (0.47) | 0.50 (0.46) | 0.60 (0.55) | 0.61 (0.58) | 0.62 (0.57) | 0.51 (0.44) | 0.60 (0.52) |

Retrieve SIF at the absorption lines

It is important to note that canopy-level radiance $(L(\lambda))$ at wavelength λ received by the spectrometer contains contributions from two sources including reflected solar energy $(r(\lambda) * \frac{I(\lambda)}{\pi})$ and upwelling SIF emission $(F(\lambda))$, that is:

$$L(\lambda) = r(\lambda) * \frac{I(\lambda)}{\pi} + F(\lambda)$$
 (1)

where $I(\lambda)$ is down-welling incoming solar irradiance; both $L(\lambda)$ and $I(\lambda)$ are provided by the SIF observation system. At some absorption lines caused by either atmosphere absorption bands or the Fraunhofer lines, however, $r(\lambda)$ and $F(\lambda)$ can be represented by some mathematical functions (Zhao et al., 2014). Within the spectral range from 680 nm to 775 nm, there are three main absorption lines including O₂-B at 687 nm, water vapor at 720 nm and O₂-A at 761 nm (Table S2).

Table S2. The central wavelengths and spectral ranges of the absorption lines used for retrieving SIF emission.

| Absorption lines | Central Wavelength (nm) | Spectral Range (nm) |
|-------------------|-------------------------|---------------------|
| O ₂ -B | 687 | 683–692 |
| Water vapor | 720 | 714–722 |
| O ₂ -A | 761 | 757–771 |

Following the Spectral Fitting Method (SFM) proposed by Zhao et al (2014), we can express $r(\lambda)$ and $F(\lambda)$ by using the second order Taylor polynomials at these three lines (λ):

$$r(\lambda) = b_1 + b_2 * (\lambda - \lambda_c) + b_3 * (\lambda - \lambda_c)^2$$
 (2)

$$F(\lambda) = b_4 + b_5 * (\lambda - \lambda_c) + b_6 * (\lambda - \lambda_c)^2$$
 (3)

where λ_c are the central wavelengths of the absorption lines (Table S2). b_1 , b_2 , b_3 , b_4 , b_5 and b_6 are six unknown coefficients in the above two equations. Combining Eq. (1), (2) and (3), we have:

$$L(\lambda) = (\lambda - \lambda_c)^2 * \frac{I(\lambda)}{\pi} * b_6 + (\lambda - \lambda_c) * \frac{I(\lambda)}{\pi} * b_5 + \frac{I(\lambda)}{\pi} * b_4 + (\lambda - \lambda_c)^2 * b_3 + (\lambda - \lambda_c) * b_2 + b_1$$
(4)

To separate $r(\lambda)$ and $F(\lambda)$ at a specific absorption line, one needs at least six measurements values of $L(\lambda)$ and $I(\lambda)$ within corresponding spectral range. After this step, we should have SIF emission at the three absorption lines named F(687), F(720) and F(761) at each instantaneous time step.

Due to poor weather conditions such as scattered clouds and heavy rainfall, some $F(\lambda)$ may have negative values which were removed from the further analysis.

References

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